



LARGE SCALE PROBLEM SOLVING USING AUTOMATIC CODE GENERATION AND DISTRIBUTED VISUALIZATION

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Abstract. Scientific computation faces multiple scalability challenges in trying to take advantage of the latest generation compute, network and graphics hardware. We present a comprehensive approach to solving four important scalability challenges: programming productivity, scalability to large numbers of processors, I/O bandwidth, and interactive visualization of large data. We describe a scenario where our integrated system is applied in the field of numerical relativity. A solver for the governing Einstein equations is generated and executed on a large computational cluster; the simulation output is distributed onto a distributed data server, and finally visualized using distributed visualization methods and high-speed networks. A demonstration of this system was awarded first place in the IEEE SCALE 2009 Challenge.

Key words: computer algebra systems, high performance computing, distributed systems, high speed networks, problem solving environment, scientific visualization, data intensive applications

1. Introduction. We describe the motivation, design, and experimental experiences of an end-to-end system for large scale, interactive and collaborative numerical simulation and visualization that addresses a set of fundamental scalability challenges for real world applications. This system was awarded first place in the IEEE SCALE 2009 Challenge in Shanghai, China in May 2009.

Our system shows a single end-to-end application capable of scaling to a large number of processors and whose output can be visualized remotely by taking advantage of high speed networking capabilities and of GPU-based parallel graphics processing resources. The four scalability challenges that are addressed are described below (see Figure 1.1).

1.1. Programming productivity. Programming productivity has long been a concern in the computational science community: the ever-growing complexity of many scientific codes make the development and maintenance of many large scale scientific applications an intimidating task. Things get even worse when one is dealing with extremely complicated systems such as the Einstein equations which, when discretized, typically result in over 20 evolved variables and thousands of source terms. In addressing these issues, we present our latest work on generic methods for generating code that solves a set of coupled nonlinear partial differential equations using the *Kranc* code generation package [1]. Our work greatly benefits from the modular design of the *Cactus* framework [2, 3], which frees domain experts from lower level programming issues, i. e., parallelism, I/O, memory management, et cetera. In this collaborative problem-solving environment based on *Cactus* and *Kranc*, application developers, either software engineers or domain experts, can contribute to a code with their expertise, thus enhancing the overall programming productivity.

1.2. Scalability to large number of processors. With the advent of Roadrunner, the first supercomputer that broke the petaflop/s mark in year 2008, the petaflop era was officially entered. There are a great number of challenges to overcome in order to fully leverage this enormous computational power to be able to solve previously unattainable scientific problems. The most urgent of all is the design and development of highly scalable and efficient scientific applications. However, the ever-growing complexity in developing such efficient parallel software always leaves a gap for many application developers to cross. We need a bridge, a computational infrastructure, which does not only hide the hardware complexity, but also provides a user friendly interface for scientific application developers to speed up scientific discoveries. In our project, we used a highly efficient computational infrastructure that is based on the *Cactus* framework and the *Carpet* AMR library [4, 5, 6].

1.3. I/O bandwidth. We are faced with difficult challenges in moving data when dealing with large datasets, challenges that arise from I/O architecture, network protocols and hardware resources: I/O archi-

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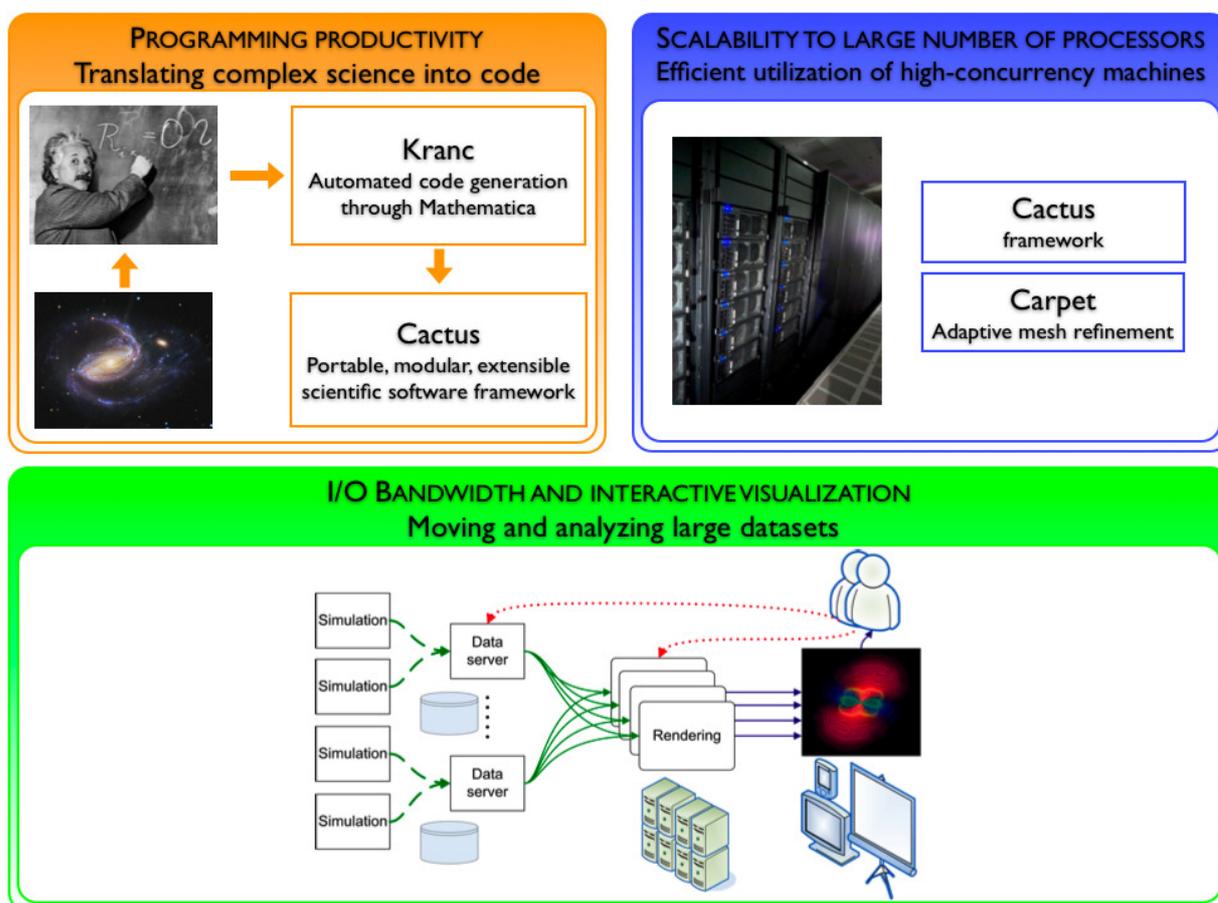


FIG. 1.1. Scalability challenges involved in an end-to-end system for large scale, interactive numerical simulation and visualization for black hole modeling.

techniques that do not use a non-blocking approach are fundamentally limiting the I/O performance; standard network protocols such as TCP cannot utilize the bandwidth available in emerging optical networks and cannot be used efficiently on wide-area networks; single disks or workstations are not able to saturate high-capacity network links. We propose a system that combines an efficient pipeline-based architecture, takes advantage of non-standard high-speed data transport protocols such as UDT, and uses distributed grid resources to increase the I/O throughput.

1.4. Interactive visualization of large data. Bringing efficient visualization and data analysis power to the end users' desktop while visualizing large data and maintain interactivity, by giving the user the ability to control and steer the visualization, is a major challenge for visualization applications today. We are looking at the case where sufficiently powerful visualization resources are not available at either the location where the data was generated or at the location where the user is visualizing it, and propose using visualization clusters in the network to interactively visualize large amounts of data.

2. Scientific Motivation: Black Holes and Gamma-Ray Bursts. Over ninety years after Einstein first proposed his theory of General Relativity, astrophysicists are increasingly interested in studying the regions of the universe where gravity is very strong and the curvature of spacetime is large.

This realm of strong curvature is notoriously difficult to investigate with conventional observational astronomy. Some phenomena might not be observable in the electromagnetic spectrum at all, and may only be visible in the gravitational spectrum, i. e., via the gravitational waves that they emit, as predicted by General Relativity. Gravitational waves have today not yet been observed directly, but have attracted great attention

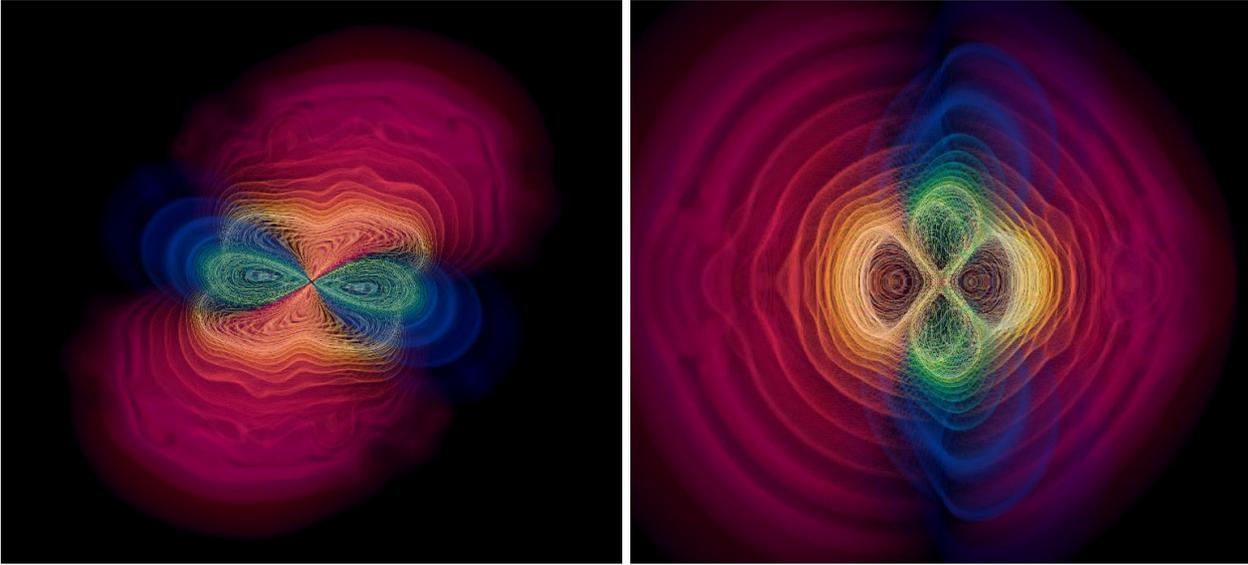


FIG. 2.1. *Volume rendering of the gravitational radiation during a binary black hole merger*

thanks to a wealth of indirect evidence [7, 8]; furthermore, gravitational wave detectors (LIGO [9], GEO [10], VIRGO [11]) will soon reach sufficient sensitivities to observe interesting astrophysical phenomena.

In order to correctly interpret the gravitational-wave astronomy data, astrophysicists must rely on computationally challenging large-scale numerical simulations to study the details of the energetic processes occurring in regions of strong curvature. Such astrophysical systems and phenomena include the birth of neutron stars or black holes in collapsing evolved massive stars, the coalescence of compact binary systems, Gamma-Ray Bursts (GRBs), active galactic nuclei harboring supermassive black holes, pulsars, and oscillating neutron stars.

Of these, Gamma-Ray Bursts (GRBs) [12] are among the most scientifically interesting. GRBs are intense, narrowly-beamed flashes of γ -rays originating at cosmological distances, and the riddle concerning their central engines and emission mechanisms is one of the most complex and challenging problems of astrophysics today. The physics necessary in such a model includes General Relativity, relativistic magneto-hydrodynamics, nuclear physics (describing nuclear reactions and the equation of state of dense matter), neutrino physics (weak interactions), and neutrino and photon radiation transport. The complexity of the GRB central engine requires a multi-physics, multi-length-scale approach that cannot be fully realized on present-day computers and will require petascale computing [13, 14].

At LSU we are performing simulations of general relativistic systems in the context of a decade-long research program in numerical relativity. One pillar of this work is focused particularly on 3D black hole physics and binary black hole inspiral and merger simulations. This includes the development of the necessary tools and techniques to carry these out, such as mesh refinement and multi-block methods, higher order numerical schemes, and formulations of the Einstein equations. A second pillar of the group's research is focused on general relativistic hydrodynamics simulations, building upon results and progress achieved with black hole system. Such simulations are crucial for detecting and interpreting signals soon expected to be recorded from ground-based laser interferometric detectors.

The specific application scenario for the work presented at the SCALE 2009 competition is the numerical modeling of the gravitational waves produced by the inspiral and merger of binary black hole systems (see Figure 2.1).

2.1. Use-case Scenario. The motivating and futuristic scenario for this work is based on enabling scientific investigation using complex application codes on very large scale compute resources:

- Scientists working together in a distributed collaboration are investigating the use of different algorithms for accurately simulating radiation transport as part of a computational model of gamma-ray bursts which uses the Cactus Framework. The resulting simulation codes use adaptive mesh refinement to dynamically add additional resolution where needed, involve hundreds of independent modules coordi-

nated by the Cactus Framework, require the use of tens of thousands of cores of modern supercomputers and take several days to complete.

- The scientists use the Kranc code generation package to automatically generate a suite of codes using the different algorithms that they wish to compare. Kranc writes these codes taking advantage of appropriate optimization strategies for the architectures on which they will be deployed, for example using GPU accelerators where available, or matching grid loops to the available cache size.
- The simulations are deployed on multiple supercomputers available to the collaboration, using co-scheduling services across different institutions to coordinate the simultaneous reservation of resources, networks, and displays. Web services are used to enable the real-time, highly configurable, collaboration of the scientists, with the simulations autonomously publishing appropriate information to services such as Twitter and Flickr.
- As the simulations run, output data is directly streamed across high speed networks to powerful GPU rendering clusters which produce the visualizations, and in turn stream their video outputs to large high resolution displays located at the collaborating sites. The displays aggregate the video outputs from each of the different simulations, allowing the scientists to visualize and compare the same output, while simultaneously interacting with and steering the visualization using tangible devices.

The scientists are thus able to use the most powerful computational resources to run the simulation and the most powerful visualization resources available to interactively visualize the data and are not limited by either their local visualization resources, or the visualization resources available at the location where the simulation is being run.

3. Automatic Parallel Code Generation.

3.1. Cactus–Carpet Computational Infrastructure. *Cactus* [2, 3] is an open source software framework consisting of a central core, the *flesh*, which connects many software components (*thorns*) through an extensible interface. *Carpet* [4, 5, 6] serves as a driver layer of the *Cactus* framework providing adaptive mesh refinement, multi-patch capability, as well as memory management, parallelism, and efficient I/O. In the *Cactus–Carpet* computational infrastructure, the simulation domain is discretized using high order finite differences on block-structured grids, employing a Berger-Oliger-style adaptive mesh refinement method [15] with sub-cycling in time, which provides both efficiency and flexibility. We use explicit Runge-Kutta methods for time integration.

Cactus is highly portable and runs on all current HPC platforms as well as on workstations and laptops on all major operating systems. Codes written using *Cactus* have been run on various brands of the fastest computers in the world, such as various Intel and AMD based systems, SGI Altix, the Japanese Earth Simulator, IBM Blue Gene, Cray XT, and the (now defunct) SiCortex architecture, among others. Recently, the *Cactus* team successfully carried out benchmark runs on 131,072 cores on the IBM Blue Gene/P at the Argonne National Laboratory [16].

3.2. Kranc Code Generation Package. *Kranc* [17, 1, 18] is a Mathematica-based computer algebra package designed to facilitate analytical manipulations of systems of tensorial equations, and to automatically generate C or Fortran code for solving initial boundary value problems. *Kranc* generates complete *Cactus* thorns, starting from a high-level description including the system of equations formulated in high-level Mathematica notation, and discretizing the equations with higher-order finite differencing. *Kranc* generated thorns make use of the Cactus Computational Toolkit, declaring to Cactus the grid functions which the simulation will use, and computing the right hand sides of the evolution equations so that the time integrator can advance the solution in time.

3.3. McLachlan Code. The *McLachlan* code [19, 20] was developed in the context of the XiRel project [21, 22], a collaboration between several numerical relativity groups worldwide to develop a highly scalable, efficient, and accurate adaptive mesh refinement layer for the *Cactus* framework, based on the *Carpet* driver, aimed at enabling numerical relativists to study the physics of black holes, neutron stars and gravitational waves. The *McLachlan* code is automatically generated using the *Kranc* code generation package (see above) from a high-level description of the underlying set of partial differential equations. The automation is of particular importance for experimenting with new formulations of the equations, new numerical methods, or particular machine specific-code optimizations. *McLachlan* employs a hybrid MPI/OpenMP parallelism.

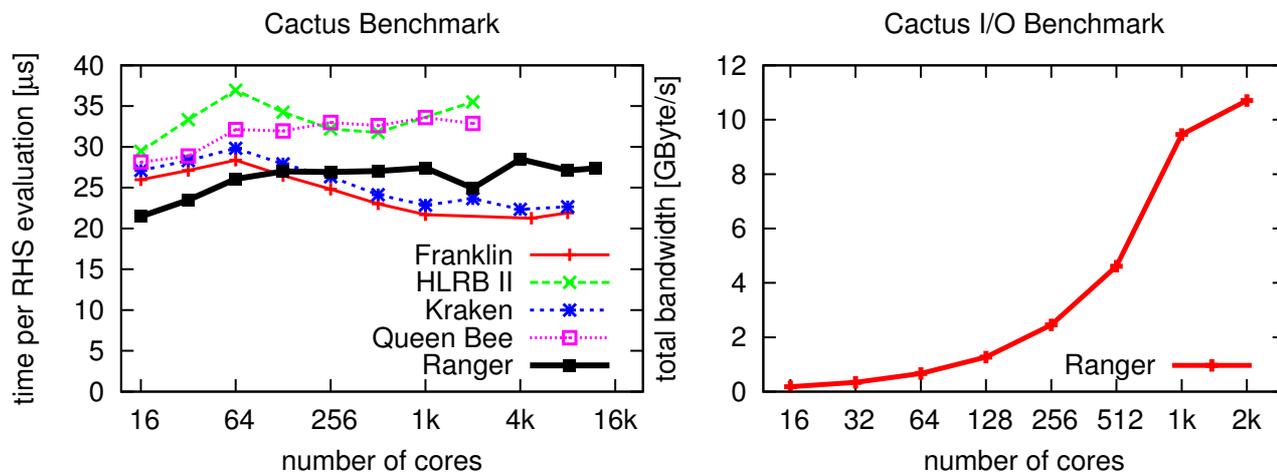


FIG. 3.1. **Left:** Weak scaling benchmark results of the McLachlan code on several current leadership HPC systems. This benchmark simulates a single black hole with nine levels of mesh refinement. The code scales well up to more than 12,000 cores of Ranger at TACC. **Right:** I/O benchmark on Ranger, showing the total I/O bandwidth vs. the number of cores. Cactus achieves a significant fraction of the maximum bandwidth already on 1,000 cores.

As can be seen from Figure 3.1, on TACC's Ranger, McLachlan and the supporting infrastructure scale well up to more than 12,000 cores. Cactus-Carpet is also able to use a significant fraction of the theoretical peak I/O bandwidth already on 1,000 cores.

4. Interactive and Collaborative Simulations.

4.1. Monitoring, Profiling and Debugging. Supporting performance and enforcing correctness of the complex, large-scale codes that Cactus generates is a non-trivial task, targeted by the NSF-funded Application-Level Performance and Correctness Analysis (Alpaca) project [23, 24].

In order to reap the benefits of the high-concurrency machines available today, it is not sufficient that a code's parallel efficiency remain constant as the size of the problem is scaled up, but also that its ease of control remains close to that of simulations carried out on a few number of computing cores; if this is not the case, the process of debugging and optimizing the code may be so time-consuming as to offset the speed-up obtained through parallelism. The Alpaca project addresses this issue through the development of application-level tools, i. e., high-level tools that are aware of the Cactus data structures and execution model.

In particular, Alpaca's objectives concentrate on three areas: (i) high-level debugging, devising debugging strategies that leverage high-level knowledge about the execution actors and the data processing, and develop tools that extract such information from a simulation and provide it in an abstract format to the user; (ii) high-level profiling, devising algorithms for extracting high-level information from timing data; and (iii) remote visualization, using visual control over the simulation data as a high-level correctness check. In particular, work within the Alpaca project includes the development of HTTPS, a Cactus module that spawns an SSL web server, with X.509 certificate authorization, at the beginning of a simulation and uses it to receive incoming connections, expose the simulation's details and provide fine-grained control over its execution.

4.2. Leveraging Web 2.0 for Collaborations. In response to the fact that computer simulations are becoming more complex and requiring more powerful resources, the way science itself is carried out is similarly changing: the growing use of computers and world-wide networks has radically modified the old custom of individual (or small-collaboration) work and hand-written data collected in notebooks. When Stephen Hawking worked out the basic theoretical framework for two colliding black holes [25] in the early seventies and Larry Smarr carried out early numerical simulations [26] a few years later, both involved only very small teams and generated perhaps a megabyte of data. The same problem has been studied in full 3D [27], now with a team size of perhaps 15 researchers, a growing number of involved institutes and an increase in generated data by a factor of about a million. Currently unsolved problems like the Gamma-Ray Burst riddle [14] will require still larger collaborations, even from different communities, and generate even more data.

In order for scientific collaborations to work at this scale, for the large amounts of data to be handled properly, and for the results to be reproducible, new methods of collaboration must be developed or already existing tools from other fields must be leveraged. Cactus can now use two tools from the latter class to announce information about simulations to existing Web 2.0 services, as described in the following.

Twitter's [28] main service is a message routing system that enables its users to send and read each others' updates, known as *tweets*. Tweets have to be very short (at most 140 characters in length) and can be sent and read via a wide range of devices, e.g. mobile texting, instant message, the web, or external applications.

Twitter provides an API [29] which allows the integration of Twitter with other web services and applications. One of the most important functions is the "statuses/update" API call, which is used to post a new Twitter message from the specified user. This Twitter API is used in a Cactus thorn to announce live information from a simulation (Figure 7.2).

Flickr [30] was launched as an image hosting website targeted at digital photographs, but short videos can be uploaded today as well. Flickr can be used at no charge with limits on the total size of images that can be uploaded (currently 100 MByte) and on the number of images which can be viewed (currently 200), along with other potential services available.

One important functionality, besides the image upload, is to be able to group images. Flickr offers a capability to group images into "Sets", and also can group different "Sets" into a "Collection". This provides a hierarchical structure for organizing simulation images.

Flickr has many features that can be taken advantage of for providing a collaborative repository for Cactus-produced images and information. All of them are accessed through a comprehensive web service API for uploading and manipulating images [31].

A new Cactus thorn uses the Flickr API to upload live images from the running simulation. Images generated by one simulation are grouped into one "Set". It is also possible to change the rendered variables, or to change the upload frequency on-line through an Cactus-internal web server (see section 4.1).

5. Distributed Visualization.

5.1. Visualization Scenario. Our scenario is the following: the visualization user is connected over a network link to a grid system of various types of resources (visualization, network, compute, data). The data to be visualized is located on a data server near the location where the scientific simulation was executed and this data server is also connected to the grid system.

There are various ways in which a visualization application can be created to solve the problem, such as running the visualization on the data server and transferring a video stream to the client, or running the visualization on the local client and transferring a data stream between the data server and the client.

These two solutions are limited by the visualization power available near the data server or near the local machine, respectively. Since powerful visualization resources are not available at the client and may not be available near the data server, we have built a three-way distributed system that uses a visualization cluster in the network, data streaming from the data server to the visualization cluster, and video streaming from the visualization cluster to the local client.

We have taken the problem one step further and considered the case where the network connection of the data server is a relatively slow one—much slower than the network capacity of the rendering machine. In this situation we are dealing with I/O scalability issues, and the solution we propose is to create a temporary distributed data server in the grid. The distributed data server uses compute and data resources that are not dedicated for this application but are allocated on-demand to support it when it needs to execute. The distributed data server can sustain much higher data transfer rates than a single data source. Data is loaded in advance from the source on the distributed data server. The architecture of this approach is illustrated in Figure 5.1. Because the visualization resources are not local to the end client, a remote interaction system is necessary in order for the user to be able to connect to and steer the visualization.

5.2. I/O. High-performance data transmission over wide-area networks is difficult to achieve. One of the main factors that can influence performance is the network transport protocol. Using unsuitable protocols on wide area network can result in bad performance—for example a few Mbit/s throughput on a 10 Gbit/s dedicated network connection using TCP. The application needs to use protocols that are suitable for the network that is utilized; our system uses the UDT [32] library for wide-area transfers in order to achieve high data transmission rates. Another issue is blocking on I/O operations: blocking I/O reduces the performance

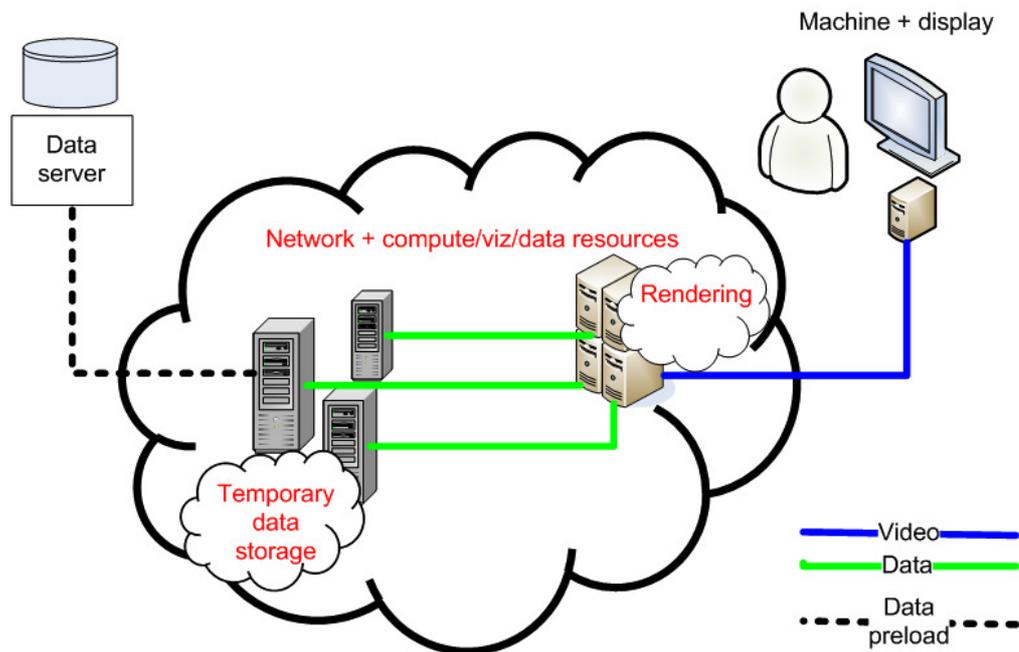


FIG. 5.1. Architecture of demonstration system which involves a temporary distributed data server allocated on-demand to improve sustained data transfer rates.

that is seen by the application, and the solution we use is based on a completely non-blocking architecture using a large number of parallel threads to keep the data flow moving.

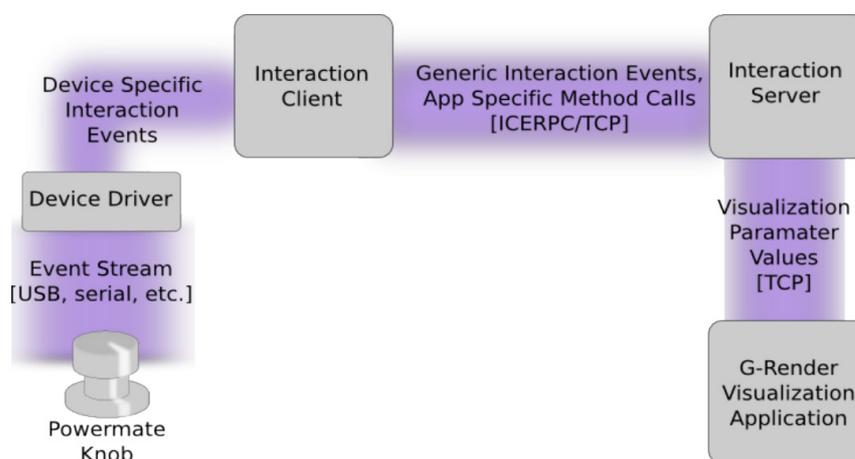
5.3. Parallel Rendering. Parallel rendering on HPC or visualization clusters is utilized to visualize large datasets. For the SCALE 2009 demonstration we have used a self-developed parallel GPU ray-casting volume rendering implementation to interactively visualize the time-dependent numerical relativity dataset, where each timestep has a size of about 1 GByte. GPU ray-casting does floating point compositing in a single pass using a fragment shader. The trade-off between rendering time and visual quality can be steered directly by the user(s). Our rendering system overlaps communication with (visual) computation, in order to achieve maximum performance. Parallelism is achieved by data domain decomposition (each node renders a distinct subsection of the data), and compositing of the resulted partial view images in order to create a single view of the entire dataset.

5.4. Video Streaming and Interaction. Interaction with the remote parallel renderer is necessary to modify navigation parameters such as the direction of viewing or the level of zoom, and to control the trade-off of visual quality and image deliver time. Since the visualization application is not local, an interaction system consisting of three components was developed. The components are a local interaction client running on the local machine, an interaction server running on the rendering cluster, and an application plug-in that connects the interaction server to the application and inserts interaction commands into the application workflow (see Fig. 5.2).

For our demonstration we used specialized interaction devices developed by the Tangible Visualization Laboratory at CCT [33] that are very useful in long-latency remote interaction systems, and can support collaboration (collaborative visualization) from multiple sites.

The final component of the system is the video streaming. Images that are generated from the remote visualization cluster need to be transported to the local client for the user to see. In the past, we have successfully utilized hardware-assisted systems running videoconferencing software (*Ultragrid* [34]) and software-based video streaming using *SAGE* [35]. Our system supports various video streaming methods including *SAGE*, a self developed video streaming subsystem, or VNC [36].

6. Related Work. OptIPuter [37] is a large project that has built an advanced infrastructure connecting computational infrastructure with high-speed “lambda” networks to create virtual distributed metacomputers. OptIPuter technologies are used in scientific applications such as microbiology [38] and climate analysis [39].

FIG. 5.2. *Architecture of interaction system*

One method for tightening the integration of applications and networks is to use reserved, exclusive access to network resources controlled by the user. Several projects, including DOE UltraScienceNet [40, 41], NSF CHEETAH [42], Phosphorus [43], G-lambda [44] and Internet2 ION [45] have explored mechanisms for providing such network services to applications.

The limitations of desktop-based visualization led to the development of parallel visualization systems and frameworks such as ImageVis3D [46], Chromium [47] and Equalizer [48] that can take advantage of computational clusters to visualize large datasets. Equalizer and Chromium are parallel rendering frameworks that can be used to build parallel rendering applications. ImageVis3D is a parallel rendering tool for interactive volume rendering of large datasets. These and other tools and techniques are being developed (for example as part of the Institute for Ultra-Scale Visualization [49, 50]) to be able to take advantage of parallel resources for visualization.

Other visualization systems such as Data Vault (DV) [51], ParaView [52], and VisIt [53] were designed to facilitate the visualization of remote data and, while they have the capability to transmit data and images over the network, they are not able to take advantage of the full capacity of high-speed networks and thus have low data transport performance, can suffer from a lack of interactivity and image quality, and do not support collaborative visualization.

Data Vault (DV) is a visualization and data analysis package for numerical codes that solve partial differential equations via grid-based methods, in particular those utilizing adaptive mesh refinement (AMR) and/or running in a parallel environment. DV provides a set of built-in functions to analyze 1D, 2D, and 3D time-dependent datasets.

ParaView is a parallel visualization application designed to handle large datasets. It supports two distribution modes: client-server and client-rendering server-data server.

VisIt is a visualization software designed to handle large datasets using client-server distribution of the visualization process. Similar to ParaView's client-server distribution, VisIt uses a parallel rendering server and a local viewer and interaction client. Most commonly, the server is as a stand-alone process that reads data from files. An alternative exists where a simulation code delivers data directly to VisIt, separating the server into two components. This allows for visualization and analysis of a live running simulation.

Several visualization systems such as RAVE [54] or the visualization system by Zhu et al. [55] are focused on the theoretical aspects of distributed visualization and do not provide the level of performance and scalability needed for current scientific applications. The distributed visualization architecture proposed by Shalf and Bethel [56] could support a variety of distributed visualization applications and inspired the development of our system.

The German TIKSL [57] and GriKSL [58] projects developed technologies for remote data access in Grid environments to provide visualization tools for numerical relativity applications. Based on GridFTP [59] and the HDF5 library [60], these projects prototyped a number of remote visualization scenarios on which this work is based.

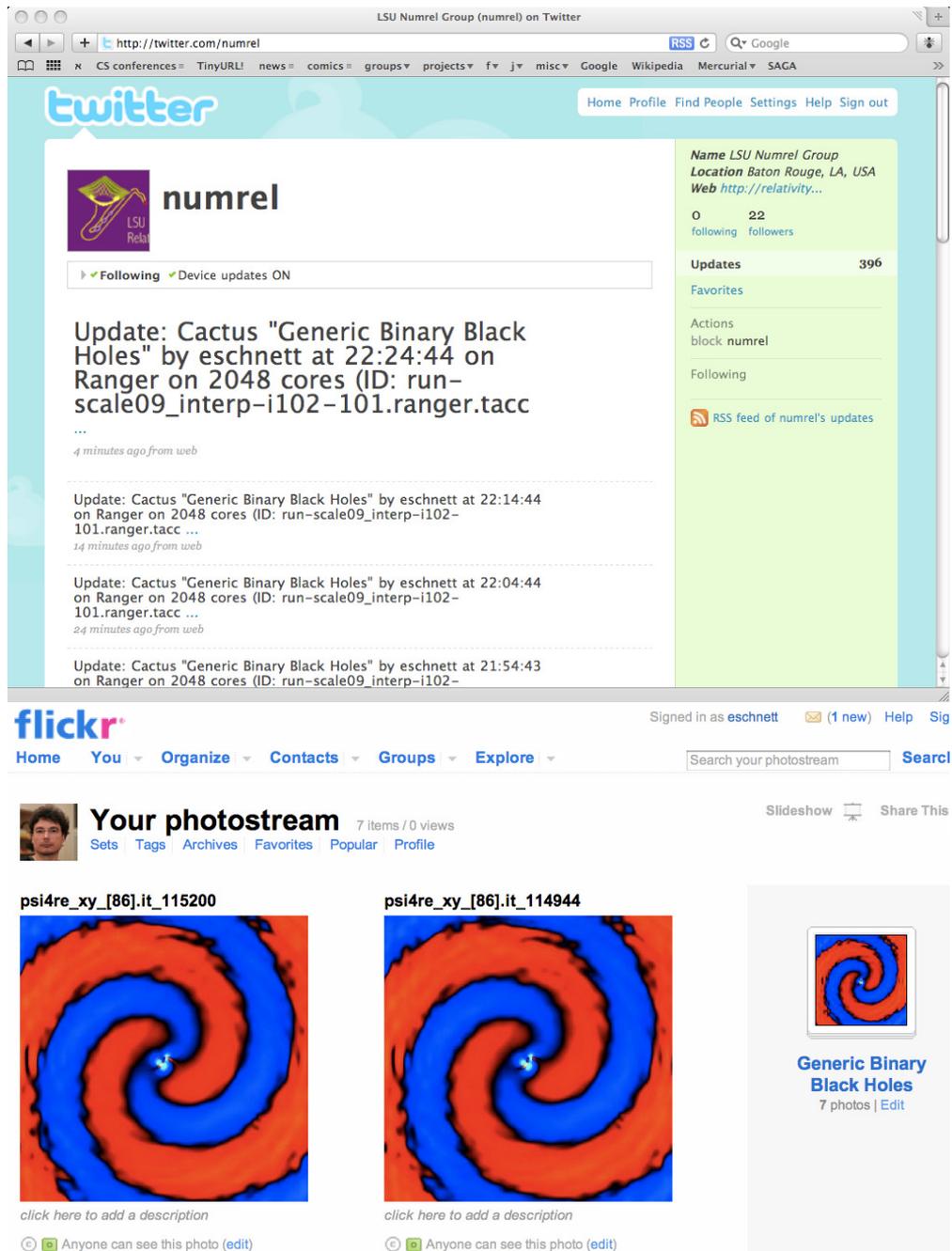


FIG. 7.2. New Cactus thorns allow simulations to announce live information and images to (top) Twitter, (bottom) Flickr enabling a new mode of scientific collaboration using Web 2.0 technologies.

8. Results After Demonstration. After SCALE 2009 we continued to improve the system, in particular the rendering and I/O subsystems. In our code generation and simulation infrastructure, we have been concentrating on adding new physics (in particular radiation transport), a new formulation of the Einstein equations, and on improving single-node (single-core and multi-core) performance.

We evaluated the visualization system performance and compared it with the performance of Paraview and VisIt using a sample dataset with a resolution of 4096^3 bytes with a total size of 64 GB.

Benchmarks were performed on the 8-node visualization cluster at LSU, each node having two Quad-core Intel Xeon E5430 processors (2.66 GHz), 16 GB RAM. The performance of the rendering cluster was improved

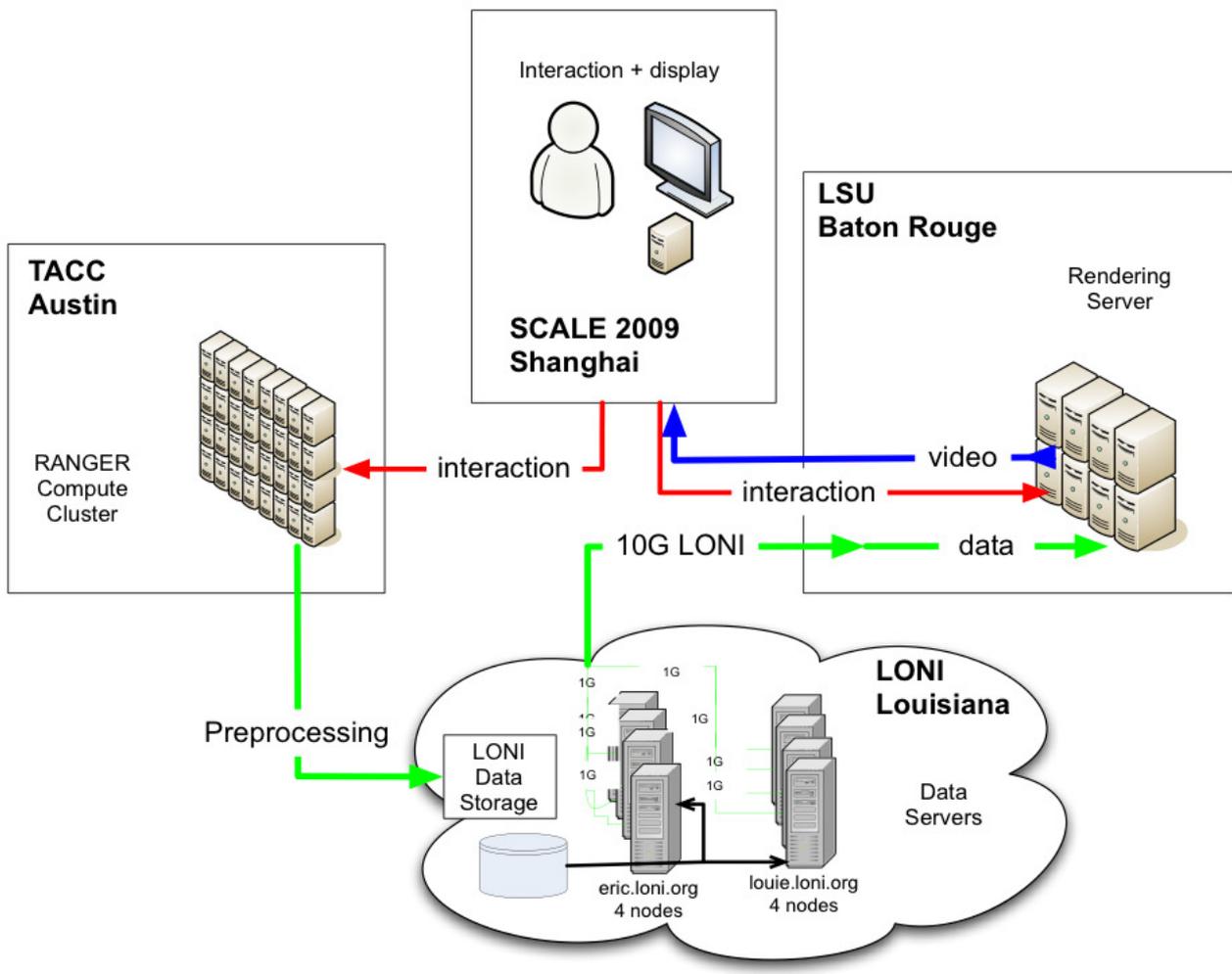


FIG. 7.3. Set up of the SCALE 2009 demonstration that involved the resources of the NSF TeraGrid, the Louisiana Optical Network Initiative (LONI) and the Center for Computation & Technology.

by upgrading the graphics hardware to four NVidia Tesla S1070 units. Each Tesla contains 4 GPUs, has 16 GB video memory and services two rendering nodes, each node thus having access to two GPU units and 8 GB video memory. The cluster interconnect is 4x Infiniband and the software was compiled and executed using MPICH2, version 1.1.1p1 using IP emulation over Infiniband.

The rendering frame rate was measured and local I/O performance was compared with remote (network) performance for three scenarios: rendering 15 GB data using 8 processes, rendering 30 GB data using 16 processes (two processes per node), and rendering 60 GB data using 32 processes (four processes per node, two processes per GPU).

The network data servers were deployed on two LONI clusters, using up to 32 distributed compute nodes to store data in the main memory. The network protocol used for data transfer was UDT.

For reference, the performance of the system when running on a single workstation was measured (workstation specifications: Intel Core2 CPU X6800, 2.93 GHz, 4 GB RAM, graphics: GeForce 8800 GTX, 1 GB video memory, 1 Gbps network interface). The rendering resolution for the benchmark is 1024x800 pixels.

The results are shown in Table 8.1. We can see that as we increase the number of rendering processes we can render more data, however the frame rate is decreasing. This reduction in speed is expected because the communication overhead increases with the number of processes. The effect is a reduction in frame rate, showing a fundamental issue with parallel rendering: at some point as the data size (and thus number of

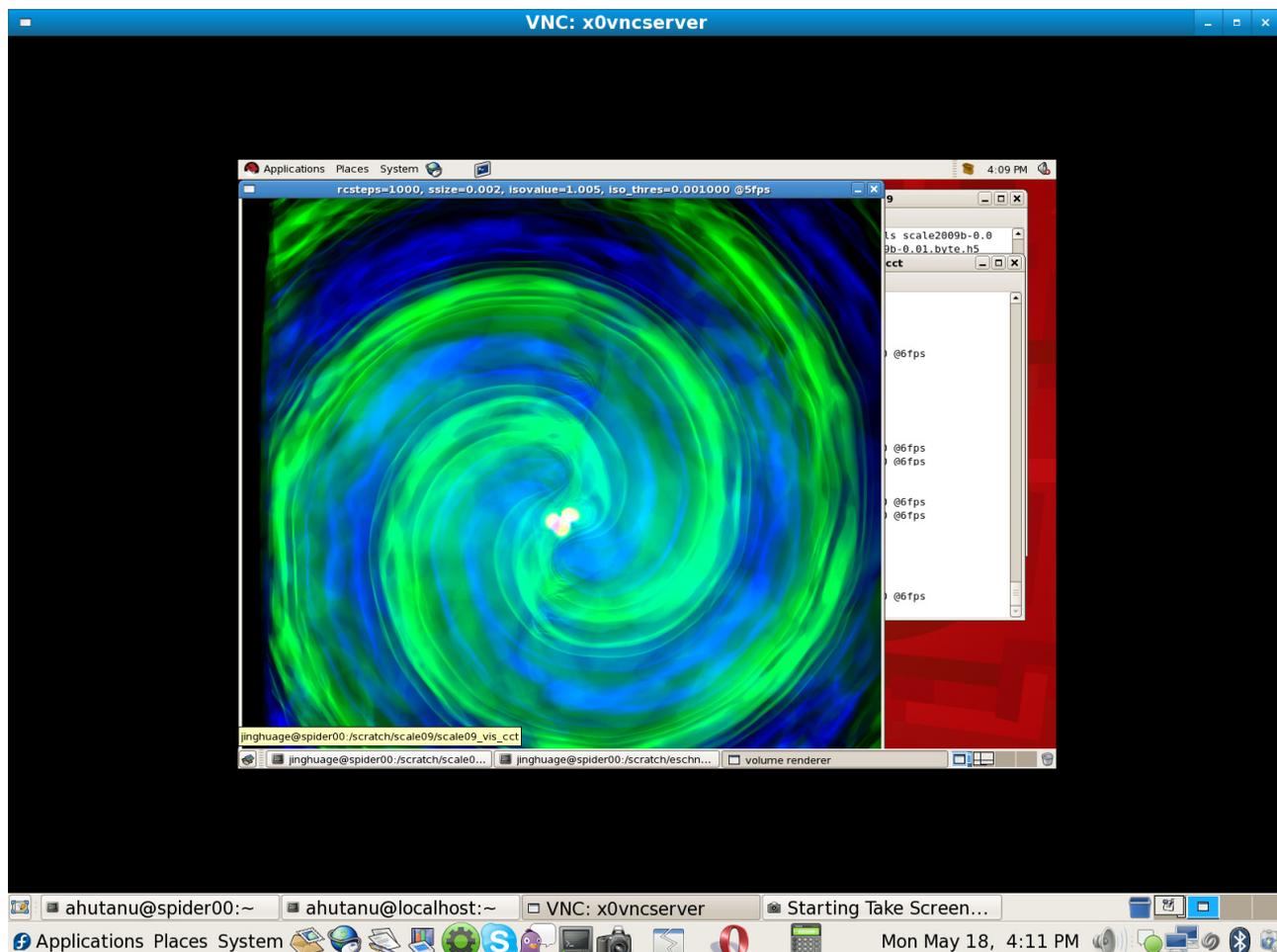


FIG. 7.4. Visualization client on the end display showing rendering of gravitational waves emitted from the inspiral collision of two black holes.

processes required to render it) increases, the frame rate drops to a level below the point of interactivity. The results show that the system is able to utilize the rendering cluster to interactively render 35 times more data than a typical workstation, and maintain an acceptable level of interactivity while rendering more than 70 times more data than on the workstation. The current system is able to interactively visualize data 60 times larger than that supported by the SCALE 2009 system.

TABLE 8.1
Data throughput and rendering scalability results.

# processes	Data size	Frame rate (fps)	Local speed	Network speed
1 (workstation)	0.8 GB	30	0.68 Gbps	0.8 Gbps
8 (cluster)	15 GB	15-21 (18 avg)	0.11 Gbps	6.6 Gbps
16 (cluster)	30 GB	11-13 (12 avg)	0.12 Gbps	5.3 Gbps
32 (cluster)	60 GB	4-5 (4.5 avg)	0.2 Gbps	4.3 Gbps

Regarding data speed, we see a big advantage when using network I/O, proving the value of the proposed approach of designing the system to be able to take advantage of high-speed networks. The system achieves 6.6 Gbps throughput over the LONI wide-area network (the limit being the network interfaces on the cluster) when using 8 processes. As we increase the number of processes the network speed decreases slightly because of the increased contention on the network interface on the same node. The remote I/O performance is 20-60

times better than local I/O performance, and the current system is able to sustain up to 6.6 Gbps transfer rates, both showing significant improvements over the SCALE 2009 system.

To better understand the features and the trade-offs of our system (named *eaviv*) a comparison with alternative visualization systems was made. Two appropriate comparison systems were identified, ParaView (version 3.6.1) and VisIt (version 1.12.0).

TABLE 8.2
Comparison of visualization systems features and performance: I/O methods, rendering and other items

Feature	<i>eaviv</i>	ParaView	VisIt
Data protocols	UDT, TCP, fully configurable	TCP only	TCP only
<i>High-speed data limit</i>	<i>Yes: Main memory</i>	<i>No: Disk size</i>	<i>No: Disk size</i>
<i>Frame rate</i>	<i>11-12 fps (30 GB)</i>	<i>0.5-1 fps (32 GB)</i>	<i>0.28-0.35 fps (32 GB)</i>
<i>Render size limit</i>	<i>60 GB (GPU memory)</i>	<i>120 GB (CPU memory)</i>	<i>120 GB (CPU memory)</i>
Collaborative support	Yes: SAGE video distribution, tangible devices	No	No
Video streaming	Parallel (SAGE)	Serial	Serial
Direct simulation connectivity	No	No	Yes
Fully-featured visualization application	No (Prototype)	Yes	Yes

The comparison was made in three different areas: data input; parallel rendering; and miscellaneous items. Both qualitative and quantitative items were analyzed. Slightly different data sizes were used due to the different modes of selecting the section of interest in each system.

Starting with data input, our system supports multiple data protocols allowing it to take advantage of high-speed networks. The benchmark results shown in Table 8.2 executed on the rendering cluster shows how our system can take advantage of the high-speed network to achieve a high data throughput. This throughput can however only be sustained for an amount of data equal to the main memory size available in the network. Both ParaView and VisIt throughput is limited by disk speed.

The second area of interest is the parallel rendering component. Our system uses a GPU-based parallel renderer, allowing it to take advantage of graphics acceleration for volume rendering and enabling high frame rate. ParaView and VisIt do not currently support parallel GPU acceleration, and in consequence the frame rate that they can achieve is below 1 frame per second. For VisIt the ray-casting parallel rendering method was used for comparison. GPU-based rendering is however limited in the data size that it can render by the amount of video memory of the graphics cards. CPU-based rendering can usually render more data, as the amount of main memory in a system is generally higher than that of video memory.

Parallel video streaming is a feature supported by our use of the SAGE system. Each rendering node, after generating a section of the final image can transmit it directly to the viewer client. In contrast, VisIt and ParaView rendering processes transmit their results first to the master node which combines them and transmits the complete image to the client. Serial video streaming introduces additional overhead and latency into the system.

Our prototype has integrated support for tangible interaction devices while allowing mouse and keyboard interaction through the use of third-party software, such as VNC. The use of SAGE and tangible interaction devices enables direct support of collaborative visualization, where multiple users, potentially at different locations around the world can simultaneously interact and collaborate using the visualization system. SAGE bridges can be used to multicast the video stream from the application to multiple users, and interaction devices deployed at each user location can be used to interact with the visualization.

One of the important missing features of our system is the lack of a direct connection between the visualization and the simulation, needed in order to visualize live data, as it is being generated (this feature is already supported by VisIt). Our system is designed as a prototype to explore the possibilities of distributed visualization, it only supports volume rendering of uniform scalar data, and has only a small fraction of the features of complete visualization systems such as VisIt and ParaView.

9. Conclusion and Future Work. Our system shows the viability of the proposed approach of using the *Cactus* framework, automated code generation, and modern numerical methods to scale to a large number of processors. It also shows how distributing the visualization system into separate components increases the amount of data that can be handled, increases the speed at which the data can be visualized compared to local techniques, and improves data transfer rates and interactivity.

The driving principle behind our approach is that, following a careful analysis and based on a detailed description of a particular technical problem, a scalable solution is to build a fully optimized integrated software system. Our proposed system is still modular, however the interfaces between the various components are flexible, and were carefully designed to enable optimizations across multiple components. The scalability of this system would suffer if each component would be designed and implemented in isolation of the other components and the coupling between components would be limited to legacy or rigid interfaces.

Our integrated approach enables us to take advantage of the latest improvements in GPU architectures, networks, innovative interaction systems and high-performance data and video transmission systems and provides a solution and a model for future scientific computing, and we believe many other applications will be able to benefit from adopting a similar approach.

In the future, we are planning to tighten the connection between the simulation and visualization components, to enable our system to visualize data on the fly as it is generated by the simulation. Our future plans include integrating automated provisioning of network resources in our application. Towards this we are currently in the process of building a testbed that connects compute, storage, graphics and display resources (provided by TeraGrid sites Louisiana State University and National Center for Supercomputing Applications and international partners such as Masaryk University) together with dynamic network circuit services provided by Internet2 and the Global Lambda Integrated Facility international community.

A fundamental service required by our system is co-allocation and advance reservation of resources. We are working together with resource providers and actively encourage them to enable these services that are crucial to enable the routine execution such complex distributed applications.

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